STScI | SPACE TELESCOPE SCIENCE INSTITUTE EXPANDING THE FRONTIERS OF SPACE ASTRONOMY

HIGH CONTRAST IMAGING (HCI) WITH JWST

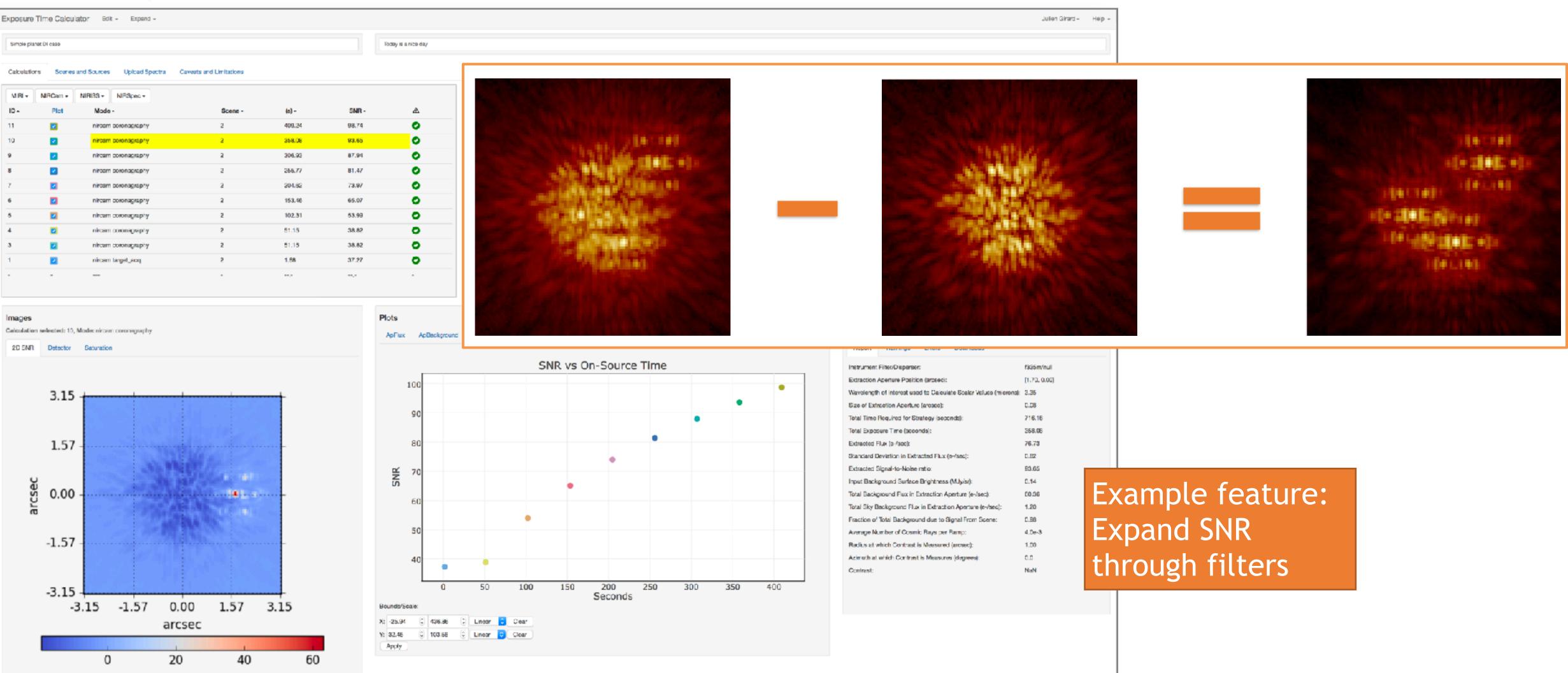
ETC Limitations & high(er) fidelity calculations: PanCAKE demo Julien Girard & Bryony Nickson

Master Class - Level 2 - Nov 19th & 20th 2019 - STScl





ETC for Coronagraphy: PSF subtraction from a reference star



Instrument Filter/Disparaer:	f335m/null
Extraction Apenture Position (arosec):	[1.70, 0.00]
Wavelength of Interest used to Calculate Scalar Values (microna):	3.35
Size of Extraction Aperture (areace):	0.08
Total Time Required for Strategy (seconds):	716.16
Total Exposure Time (accords):	358.08
Extracted Flux (o-/acc):	76.73
Standard Deviction in Extracted Flux (e-/sec):	0.82
Extracted Signal-to-Noise ratio:	83.65
Input Background Surface Brightness (MJy/sr):	0.14
Total Background Flux in Extraction Aperture (e-/sec):	00.00
Total Sky Background Flux in Extraction Aperture (e-/sec):	1.20
Fraction of Total Background due to Signal From Scene:	0.98
Average Number of Cosmic Rays per Bamp:	4.0e-3
Radius at which Contrast is Measured (arcsec):	1.00
Azimuth at which Contrast is Measures (degrees):	0.0
Contrast:	NaN

<u>jwst.etc.stsci.edu</u>



ETC for Coronagraphy: Limitations for High Contrast Imaging

Pre-computed PSF library from WebbPSF with a discrete number of angular separations (sparse spatial sampling)

→ Calculations can be inaccurate in the speckle limited regime (close to the coronagraphs, typically at separations < 1")

ETC does not account for spectral mismatch (only photometrically) of the PSF reference star ETC supposes a perfect centering (target acquisition) of all stars

→ Calculations can be optimistic

PSF calculations "on the fly" are time consuming: can be done in command line with Pandeia engine or with PanCAKE (not yet fully supported)

- Custom small grid dithers and positioning
- Custom spectral sampling
- Custom field of view

The ETC PSF subtraction strategies assume the same detector readout parameters for all stars in a workbook → If one wants to use a brighter reference star, several ETC workbooks are needed

The ETC cannot inject ring like features or disks

Can be done with pyNRC (not supported by STScI)

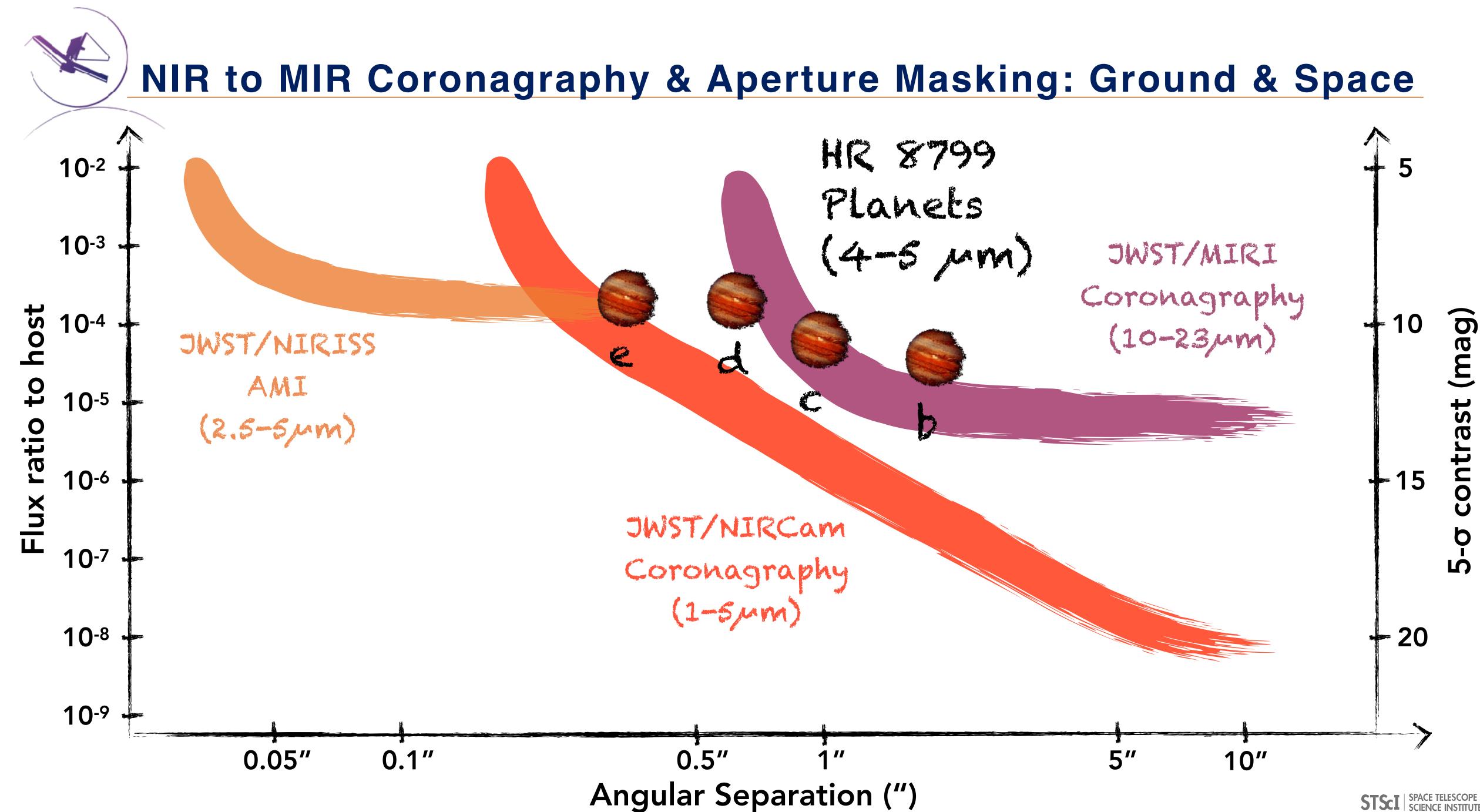
https://github.com/spacetelescope/ pandeia-coronagraphy

https://pynrc.readthedocs.io













Factors degrading PSF calibration and subtraction

Wavefront drifts of the observatory

PSF star color differences

Self-subtraction biases (esp. for disks)

Imperfect target acquisitions

Line-of-sight jitter and dynamic wavefront error

Optional:

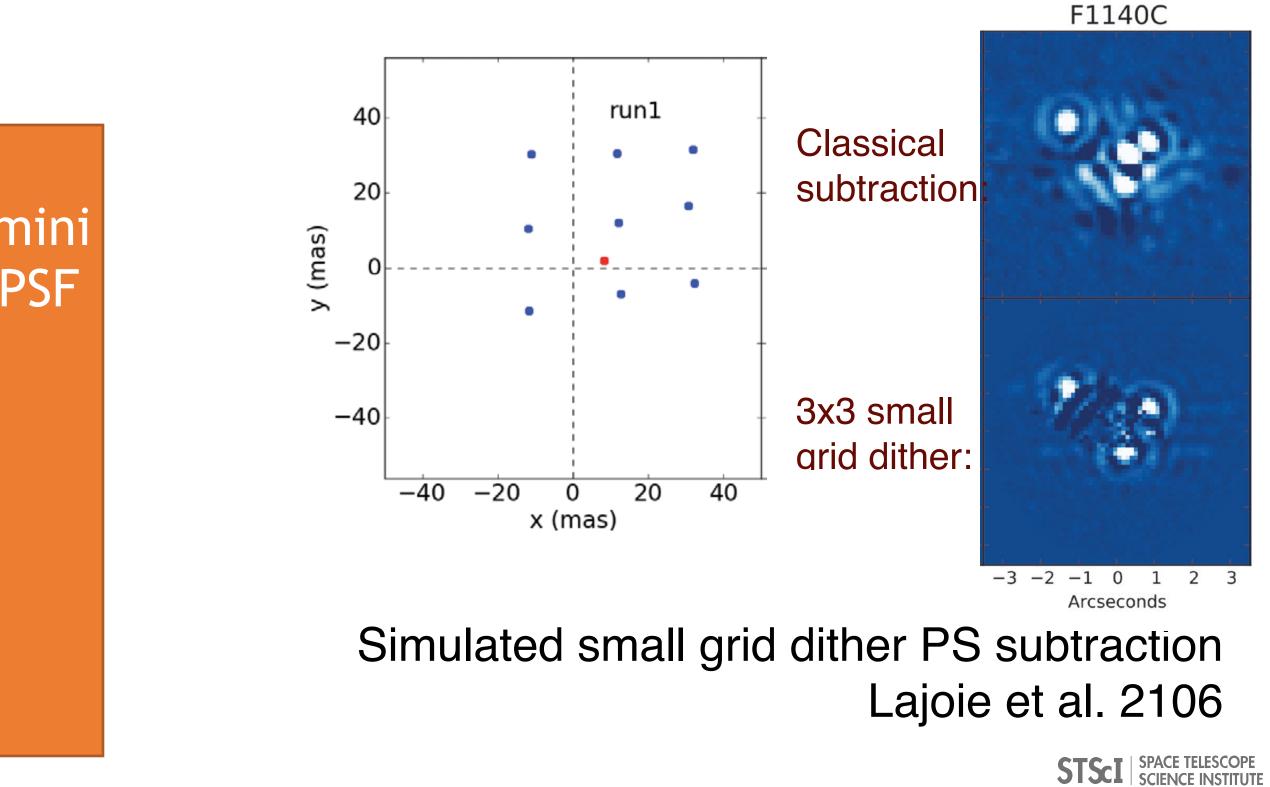
Perform sub pixel dithers of the PSF star to build mini PSF reference library, then synthesize an optimal PSF that matches the target position precisely. "Small Grid Dithers".

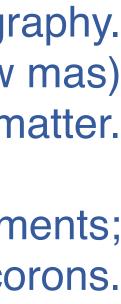
>10x contrast improvement for MIRI, 3-5x improvement for NIRCam.

Cost: 5-9x longer PSF star exposure times.

Target acq is required for all coronagraphy. Expected precision is very good (~ few mas) but residuals still matter.

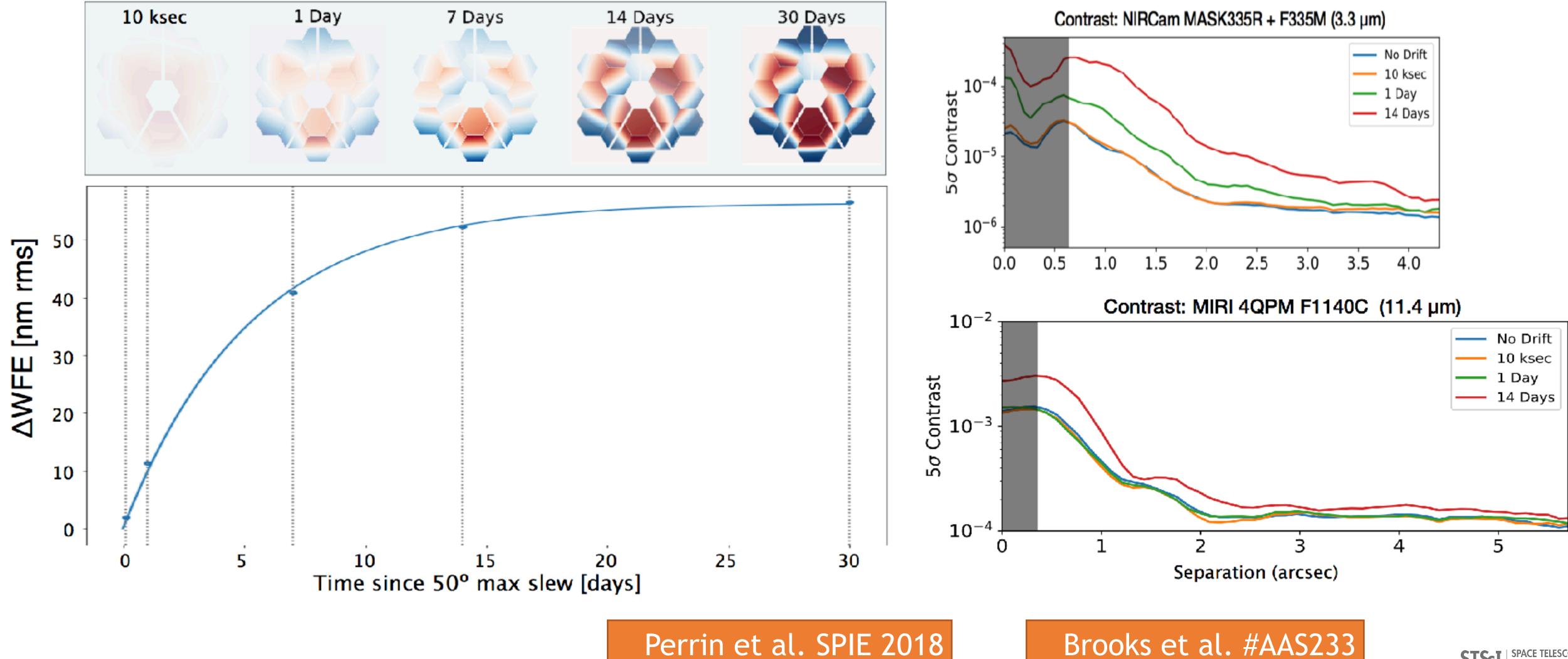
All coronagraphs are sensitive to misalignments; MIRI FQPMs more sensitive than NIRCam Lyot corons.



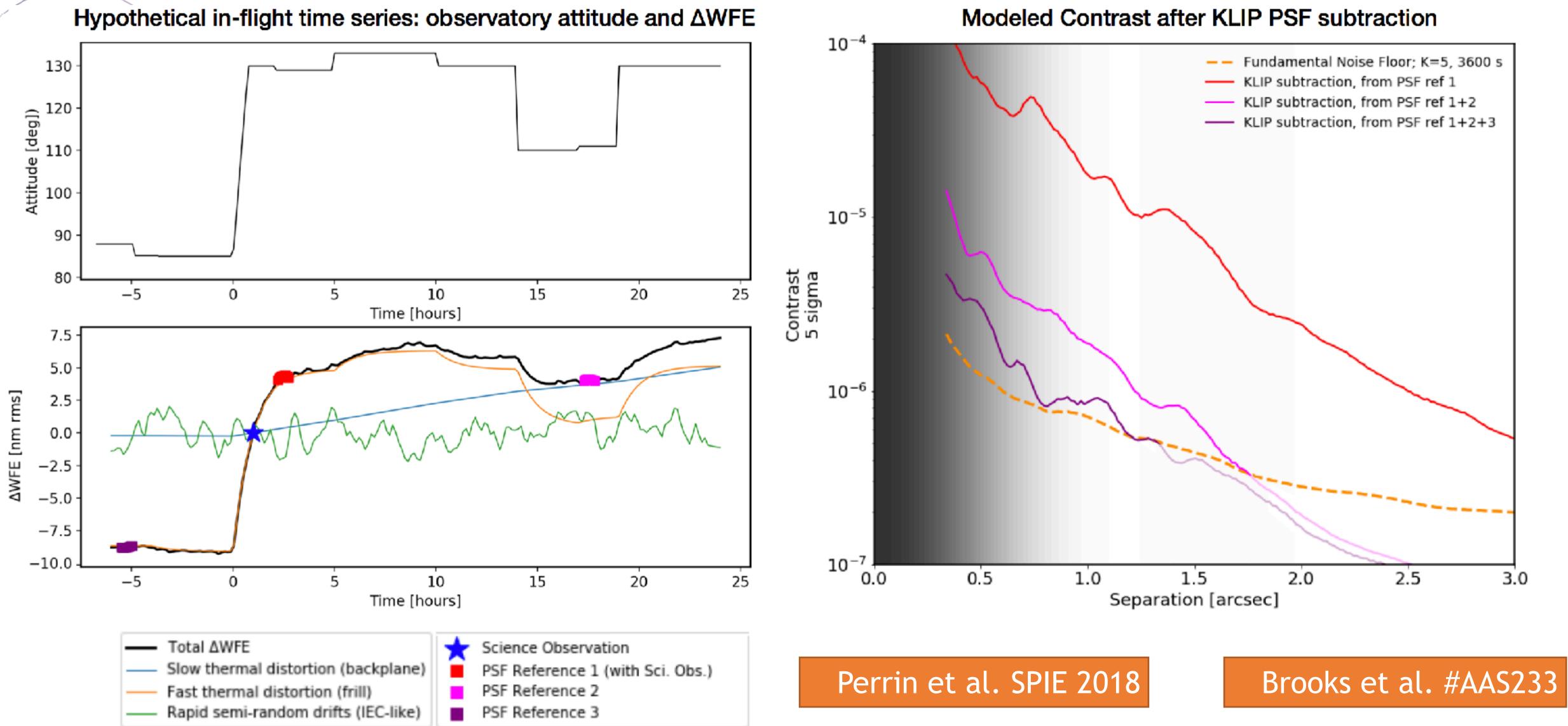


Towards the generation of realistic datasets : impact of slew

Model-predicted ΔWFE from maximum hot-to-cold slew





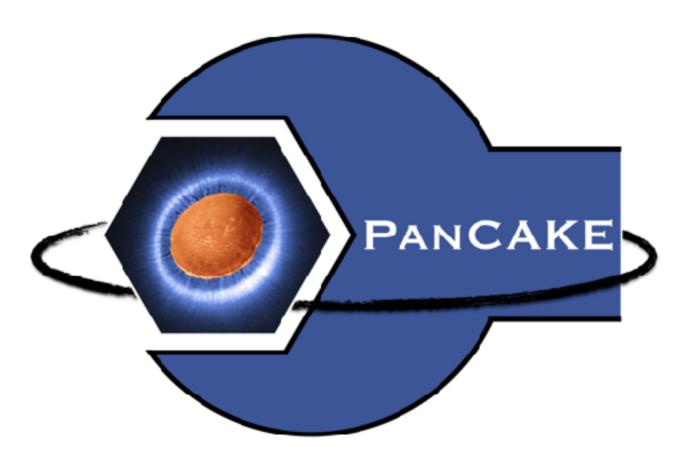


Towards the generation of realistic datasets : impact of slew

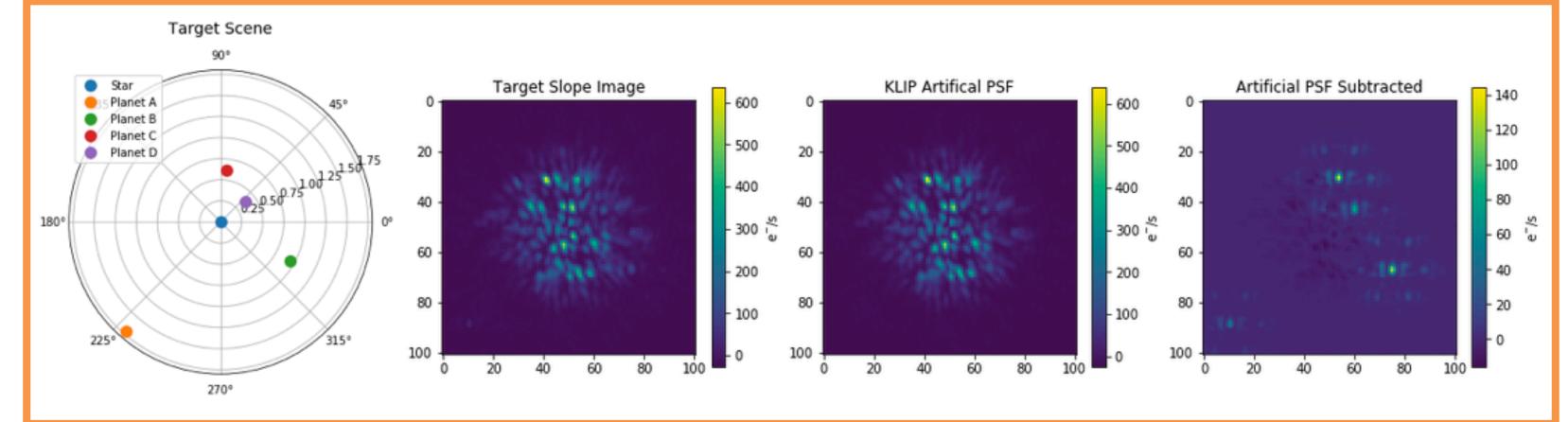




<u>github.com/spacetelescope/pandeia-coronagraphy</u>



Can call WebbPSF "on the fly" Contrast curves

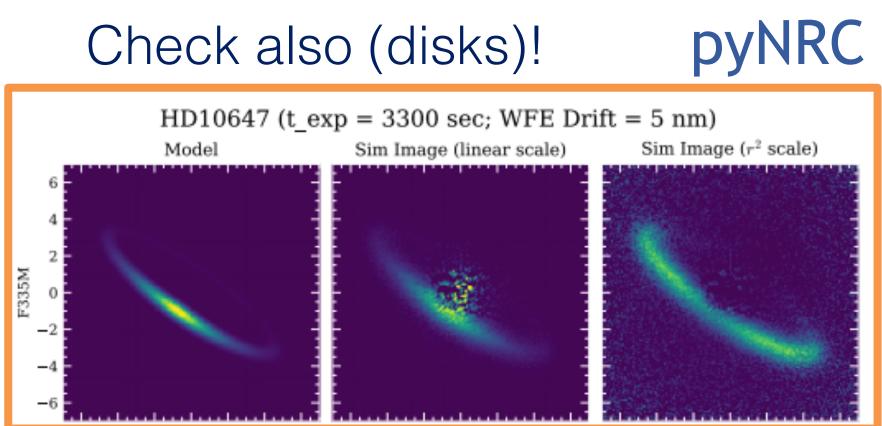


Van Gorkom, York, Perrin, Girard



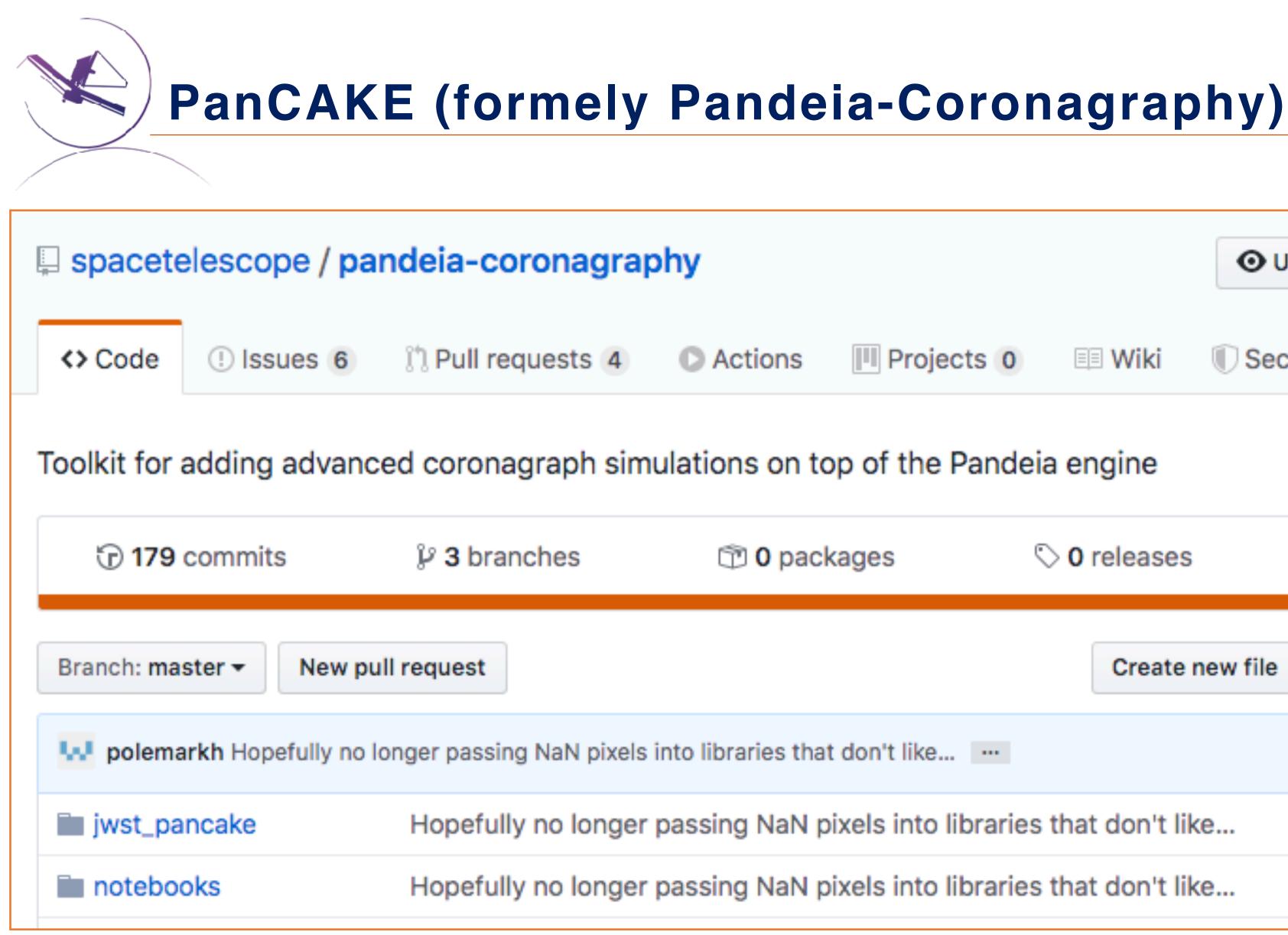
custom PSF grids/dithers, FoV, spectral sampling, more precise extractions Several coronagraphic specific functions

Currently improving the scene compatibility with ETC UI and outputs



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https://github.com/spacetelescope/pandeia-coronagraphy

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Selecting a PSF Reference Star (avoid spectral mismatch)

					A	lver	age	Sei	nsiti	vity	Los	ss at	t r =	0.5	-2''	(F	200	W+1	MAS	SK2	10R	.)					1
Μ	5V	1.17	1.17	1.16	1.15	1.14	1.14	1.13	1.13	1.12	1.12	1.11	1.11	1.10	1.06	1.00	0.83	0.48	0.02	0.00	1.11	1.09	1.08	1.05	0.99		1
Μ	2V	1.12	1.11	1.10	1.10	1.08	1.08	1.07	1.06	1.06	1.06	1.05	1.05	1.04	1.00	0.93	0.75	0.39	0.00	0.02	1.05	1.03	1.02	0.99	0.92		
Μ	0V	0.69	0.68	0.66	0.66	0.64	0.64	0.63	0.62	0.61	0.60	0.60	0.59	0.58	0.53	0.44	0.23	0.00	0.40	0.50	0.60	0.58	0.55	0.50	0.42		
K	7V	0.32	0.32	0.30	0.30	0.28	0.27	0.26	0.25	0.24	0.23	0.23	0.22	0.21	0.16	0.09	0.00	0.23	0.75	0.82	0.23	0.20	0.18	0.15	0.08		
K	5V	0.12	0.12	0.11	0.10	0.09	0.08	0.07	0.07	0.06	0.06	0.05	0.05	0.04	0.02	0.00	0.08	0.45	0.92	0.99	0.05	0.04	0.03	0.01	0.01		
φĸ	2V	0.06	0.06	0.05	0.05	0.03	0.03	0.03	0.02	0.01	0.01	0.01	0.01	0.01	0.00	0.01	0.16	0.53	0.99	1.05	0.01	0.01	0.00	0.01	0.03		
<u>F</u> K	.0V	0.03	0.03	0.02	0.02	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.04	0.21	0.58	1.03	1.09	0.00	0.00	0.00	0.02	0.05		
G G G	8V	0.02	0.02	0.02	0.02	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.04	0.22	0.59	1.04	1.10	0.00	0.00	0.00	0.02	0.06		
G et	5V	0.02	0.02	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.05	0.23	0.60	1.04	1.10	0.00	0.00	0.01	0.02	0.06		
g G	2V	0.02	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.05	0.24	0.61	1.05	1.11	0.00	0.00	0.01	0.02	0.07		
G	0V	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.06	0.24	0.62	1.05	1.11	0.00	0.00	0.01	0.02	0.07		
Эц F	8V	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.06	0.25	0.62	1.06	1.12	0.00	0.00	0.01	0.03	0.08		0.4
F CIG	5V	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.07	0.26	0.63	1.06	1.12	0.00	0.00	0.01	0.03	0.08		
ר _F	2V	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.03	0.08	0.27	0.64	1.07	1.13	0.00	0.01	0.02	0.04	0.09		
F	0V	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.03	0.08	0.28	0.65	1.08	1.14	0.01	0.01	0.02	0.04	0.10		
А	5V	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.04	0.10	0.30	0.67	1.09	1.15	0.01	0.02	0.03	0.06	0.11		
А	.3V	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.02	0.04	0.10	0.30	0.67	1.09	1.15	0.01	0.02	0.03	0.05	0.11		
А	1V	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.05	0.12	0.32	0.68	1.11	1.16	0.02	0.03	0.04	0.07	0.14		
А	.0V	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.03	0.06	0.12	0.32	0.69	1.11	1.17	0.02	0.03	0.04	0.07	0.14		
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Reference Star SpType

Credit: Jarron Leisenring (pyNRC)

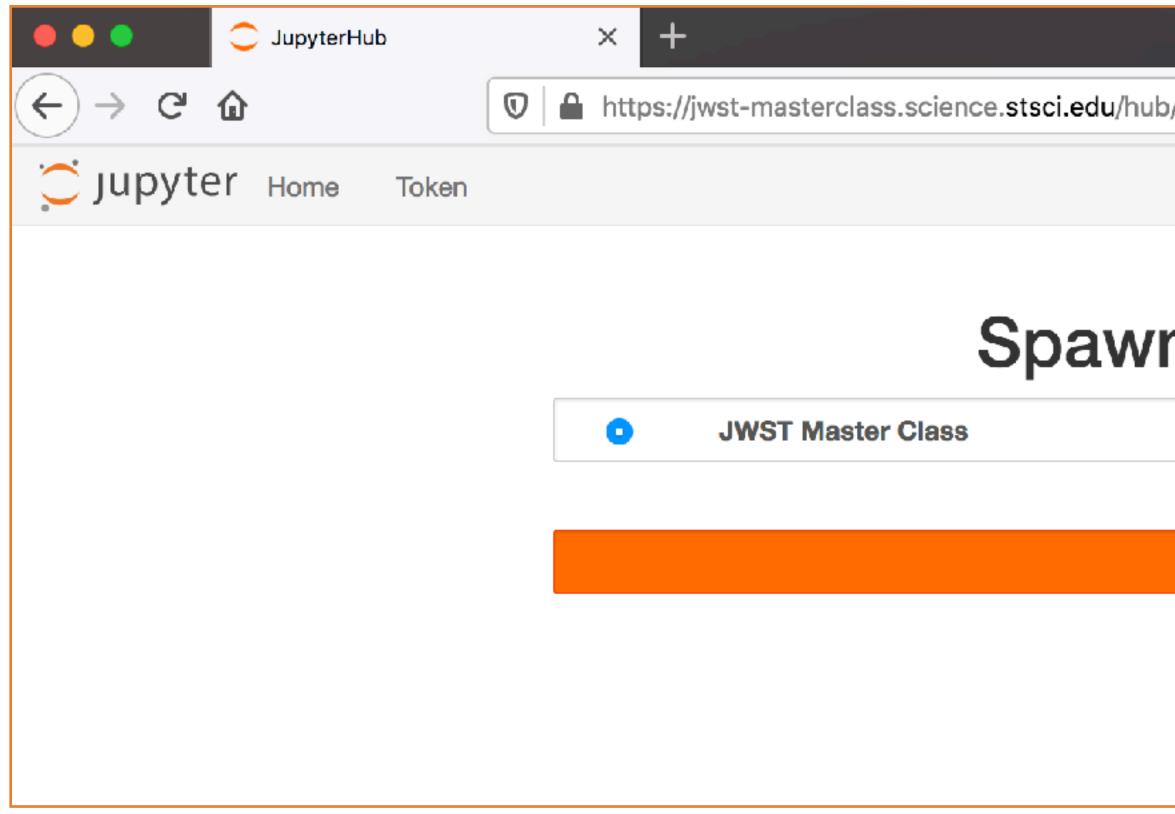
More critical at shorter wavelengths (NIRCam), not critical for MIRI







<u>https://jwst-masterclass.science.stsci.edu</u>



May take a few minutes to load

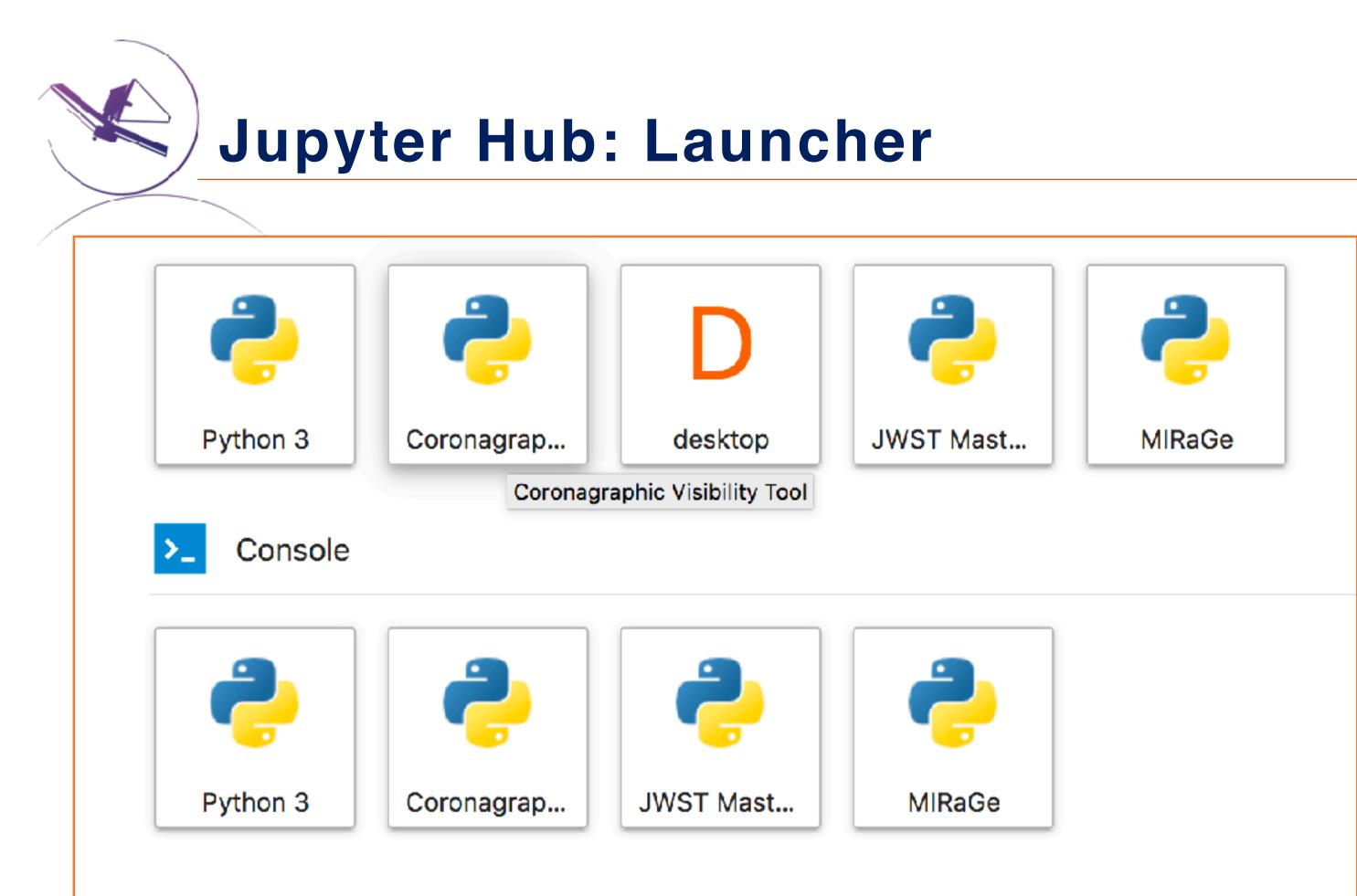
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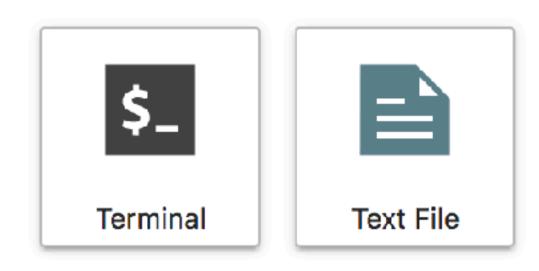
Spawn







Other



https://jwstmasterclass.science <u>stsci.edu</u>



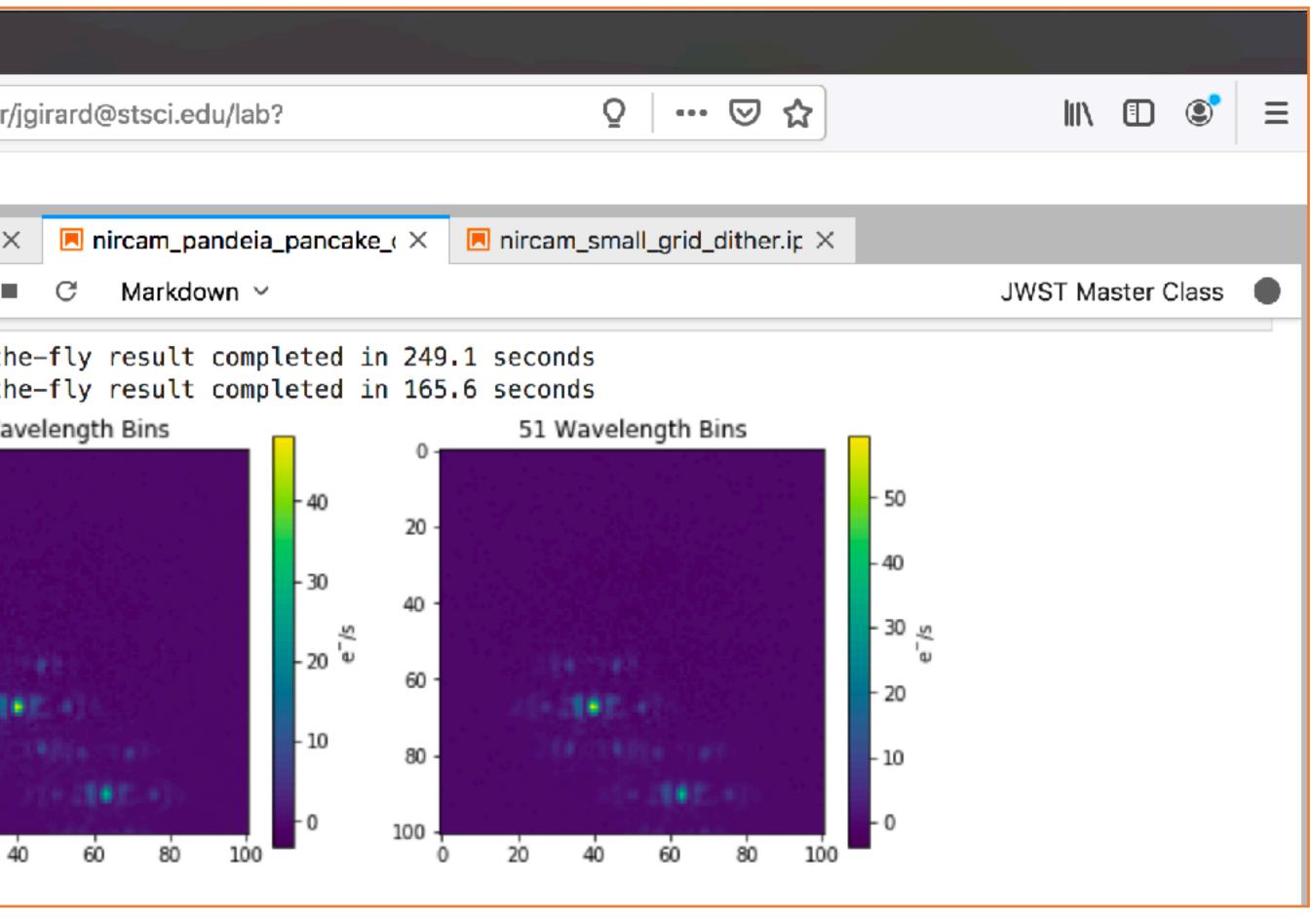
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Jupyter Hub: "on-the-fly" calculations with PanCAKE

Pandeia/WebbPSF & Small Grid dithers (SGD)

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NIRCam and MIRI Coronagraphy of HR 8799 bcde

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Jupyter Hub: JWST Master Class PanCAKE notebook

