STSCI | SPACE TELESCOPE SCIENCE INSTITUTE **EXPANDING THE FRONTIERS OF SPACE ASTRONOMY**

HIGH CONTRAST IMAGING (HCI) WITH JWST

- - Master Class Level 2 Nov 19th & 20th 2019 STScl

General & Coronagraphy Julien Girard, Bryony Nickson, Tony Roman, Beth Perriello, Bill Blair

Hands-on Session









- JMMC's SearchCal
- **Google Search "JMMC SearchCal"**
- http://www.jmmc.fr/searchcal page.htm Or
- The Coronagraph Visibility Tool (CVT): **Google Search "JWST CVT", GitHUB** Or https://jwst-docs.stsci.edu/other-tools/target-

github.com/spacetelescope/ jwst coronagraph visibility

visibility-tools/jwst-coronagraphic-visibility-tool-help







and shine

Use Case: HR 8799 b c d e



The HR 8799 Science Use Case

Face on

- 4 planets with mass $< 8 M_{Jup}$
- b at 1.7" is the faintest
- b c d e are all doable with NIRCam Coronagraphy
- b & c can be done with MIRI 4QFM coronagraph
- e can be attempted with NIRISS/AMI













and suit

Reference Star Selection





Find a suitable brighter reference star (3 < K < 5) within 5° of HR8799

(with a similar spectral type if possible)

- **Using Simbad**
- Using JMMC/SearchCal \blacklozenge
- Using the USNO single star catalog •
- Using Python astropy, astroquery •

Be careful!!

- ***** More than one solution
- for PSF subtraction or as AMI calibrator!
- HST or even 2MASS diffraction spikes

Exercise: Search for a more efficient reference star, methods

* Mind the risks: a close binary (with separation 0.1" to 2") will be useless * As much as possible, check archival high spatial resolution images (AO,



Keep in mind: slew & guide star overheads

Changing attitude

- 1. Update observatory pointing and roll
- 2. Let disturbances settle
- 3. Reacquire guide star
 - + Fine guide (always)
 - +Track (>0.06")
 - +Acquisition (>25")
 - Identification (new visit)



Simbad search by coordinates and classify by K mag

N	Identifier	dist(asec)	Otype	ICRS (J2000) RA	ICRS (J2000) DEC	Mag U	Mag B	Mag V	Mag R	Mag I	Mag H	Mag K	Sp type	#re 200 202
$ \Delta \nabla$			$\Delta \nabla$			$\Delta \nabla$	$\Delta \nabla$	$ \Delta \nabla $	$\Delta \nabla$	$\Delta \nabla$	$\Delta \nabla$			
4955	HD 216633	11963.06	*	22 54 14.8500913241	+22 24 07.145994558		9.74	8.37			5.377	5.19	K5	
9321	BD+23 4733	16012.33	*	23 21 39.7488092749	+24 08 49.513412865		11.36	9.61			5.436	5.196	K7	
1215	HD 218663	5852.94	*	23 09 40.2124602869	+19 35 30.128156497		10.33	8.76			5.405	5.209	K2	
175	BD+20 5273	1866.64	*	23 05 51.7949829456	+21 29 27.828869410		10.82	9.05			5.437	5.224	K5	
5	HD 218396	0.00	El*	23 07 28.7156905667	+21 08 03.302133882		6.21	5.953			5.280	5.240	F0+VkA5mA5	79
6801	IRAS 23161+1809	13585.27	*	23 18 39.4349900926	+18 25 38.659426227						5.63	5.245	~	
6885	HD 219446	13655.09	*	23 15 41.0513245768	+24 25 20.455058608		8.85	7.80			5.377	5.269	G9III	
3613	BD+22 4761	10298.69	*	23 02 13.8255473763	+23 43 30.995468478		10.03	8.59			5.470	5.294	K5	
9136	BD+22 4824	15845.62	*	23 24 09.7717717457	+23 14 47.010612289		10.32	8.82			5.491	5.294	K7	
2507	IRAS 23088+1841	8458.76	*	23 11 21.3650313391	+18 58 05.497297693		11.64	10.29			5.591	5.295	~	
8510	BD+23 4641	15188.37	*	22 57 05.5881313232	+24 36 35.558634258		10.97	9.32			5.548	5.310	K7	

HR 8799 at 0" (center of the search)

http://simbad.harvard.edu/simbad/sim-coo









N	Identifier	dist(asec)	Otype	ICRS (J2000) RA	ICRS (J2000) DEC	Mag U	Mag B	Mag V	Mag R	Mag I	Mag H	Mag K	Sp type	#re 200 202
$\Box \nabla$	$\Delta \nabla$	$\Delta \nabla$	$\Box \nabla$	$\Delta \nabla$	$\Delta \nabla$	$ \Delta \nabla $	$\Box \nabla$	$\Box \nabla$	$\Box \nabla$	$ \Delta \nabla $	$\Delta \nabla$	$ \Delta \mathbf{\nabla} $	$\Delta \nabla$	
8910	* 56 Peg	15605.37	SB*	23 07 06.7390851	+25 28 05.732922	7.20	6.06	4.74	3.74	3.07	1.830	1.79	K0.5II:Ba1CN- 2CH-0.5	e
4947	V* GO Peg	11946.06	LP*	22 55 00.9741276721	+19 33 35.024093025		9.08	7.45			2.214	1.842	~	
9104	V* V336 Peg	15812.40	LP*	22 54 40.3656099230	+24 23 13.703275292		8.98	7.48			2.068	1.888	M4III	
2237	V* V338 Peg	7974.51	LP?	22 58 06.4171887956	+21 30 47.434833192		8.84	7.38			2.294	1.942	~	
7182	V* BI Peg	13871.14	Mi*	22 57 51.7690882945	+18 01 00.797492648		12.47	10.80			2.583	2.124	M9	
5159	HD 220211	12191.55	۷*	23 21 49.2694060590	+20 38 15.939509092		8.89	7.16			2.400	2.156	~	
2560	BD+18 5085	8572.37	*	23 00 29.9457750451	+19 24 16.958016197		11.26	9.90			3.009	2.597	M2	
6163	HD 218792	13035.57	*	23 10 42.6364554967	+17 35 39.741311008		7.012	5.661			2.889	2.656	K4III	1
2862	BD+22 4768	9113.89	*	23 03 41.3486608588	+23 30 33.831124671		10.61	8.85			3.691	3.217	MO	
8662	BD+24 4689	15339.38	*	22 58 59.7609593773	+24 55 20.061833612		10.20	8.59			3.567	3.227	M2	
5774	HD 219992	12728.31	۷*	23 20 11.6244960670	+23 05 28.155121788		8.03	6.62			3.590	3.349	K2	
5954	HD 216786	12877.52	V*	22 55 40.2651162827	+18 52 30.707097914		9.61	8.04			3.726	3.437	~	
1894	HD 219525	7335,45	*	23 16 12.4872884220	+21 03 20.847662279		9.95	8.23			3.985	3.699	K5	
2535	* 51 Peg	8517.35	**	22 57 27.9804167474	+20 46 07.782240714	6.39	6.16	5.46	4.97	4.61	4.23	3.911	G2IV	69
1615	BD+22 4781	6859.62	*	23 08 18.1092867832	+23 01 48.480462621		10.21	8.45			4.417	3.982	MO	
1536	HD 217636	6621.66	*	23 01 54.4372679841	+19 50 15.814475079		8.58	7.16			4.171	3.990	K2	
2084	HD 219196	7703.31	*	23 13 59.3604593407	+19 38 02.045114663		7.65	6.47			4.100	4.015	K2	
9234	HD 216201	15924.61	PM*	22 50 39.1282094833	+19 08 28.342855350	8.610	7.600	6.490			4.447	4.031	КО	

51 Peg at 2.4° seems the best one (G type, K = 3.9) We know it's not as binary (no problem for the PSF subtraction)

Simbad: look for a K< 5 star with a distance < 18,000 arcsec





JMMC/SearchCal: developed for long-baseline interferometry

	e e SearchCal [c1]												
-Query P	Parameters												
-1) Ins	trumental Configurat	tion	2) Science	Object –				1	3) SearchCal Parameters				
	Magnitude Ban	nd : K			Na	me : 🔍 HR 87	799	8	Min. Magnitude (K) : 3.24				
	Wavelength (K) [un	n] · 2.2		RA 2000 [hh:mm:ss] : 23 07 28 7156905667				Max, Magnitude (K): 7.24					
				107 20		35]. <u>250720</u> .	1110505007			j			
	Max. Baseline [n	n]: <u>102.45</u>	DE	C 2000	[+/-dd:mm:	ss]: +21 08 0	3.302133882			Scena	ario :	🗿 Bright	🔵 Faint
					Magnitude	(K): 5.24				RA Range [r	nn] : 240.	.0	
									D	EC Bancia (d	$\frac{1}{200}$)	
									D	EC Range lu	eg]. 50.0	,	
Progr	ess :											Get	Calibrators
Found (Calibrators (4959 sou	urces. 4804 filtere	ed)										
Index	dist A HD	RAI2000	DEI2000	К	SpType	SIMBAD	ObiTypes	vis2	vis2Err	diam chi2	LDD	e LDD rel	
1	0.016	23 07 24.5827	+21 08 01.0608	6.843	G0	BD+20 5278p	,Star,*,IR,	0.99	2.396E-4	2.874	0.199	2.466	0.187
2	0.519 HD identifier	, click to call Simbad	on this object 7.8412	5.224	К5	BD+20 5273	,Star,*,IR,	0.969	0.001	0.102	0.513	2.272	0.473
3	0.553 218381	23 07 25.5058	+20 34 53.8644	4.414	К0	HD 218381	. ,Star,**,*,IR,	, 0.952	2 0.002	0.004	0.641	2.39	0.597
4	0.709	23 06 40.7455	+21 49 07.0176	5.579	К0	HD 218302	,Star,*,IR,	, 0.982	8.226E-4	0.453	0.391	2.216	0.365
5	0.738	23 10 38.3014	+21 10 59.5272	5.154	К2	<u>HD 218791</u>	,Star,IR,*,	, 0.97	0.001	0.83	0.5	2.242	0.463
6	0.841	23 03 58.3272	+20 56 28.1616	6.713	G/K	<u>EYC</u>	,PM*,PM*,*,*,I	. 0.993	7.865E-4	1	0.241	5.542	0.223
7	0.997 <u>218172</u>	23 05 35.3270	+20 14 27.3588	5.85	F8IV	<u>HD 218172</u>	,Star,*,IR,	, 0.99	€ 4.792E-4	0.327	0.287	2.378	0.271
8	1.193	23 11 30.2443	+21 52 22.7640	6.824	К5	HD 218895	,Star,*,IR,	, 0.993	3.072E-4	2.373	0.235	2.28	0.217
9	1.445	23 01 17.3088	+21 12 58.7124	5.426	К2	HD 217557	,Star,*,IR,	, 0.977	0.001	0.12	0.438	2.206	0.407
10	1.48	23 02 05.6947	+20 21 21.4776	4.785	К2	HD 217660	,Star,IR,*,	0.963	0.002	0.77	0.577	2.518	0.535
11	1.515	23 01 02.6974	+21 21 02.4516	6.106	К0	BD+20 5258	,Star,*,IR,	, 0.988	3 5.224E-4	0.282	0.311	2.216	0.29
12	1.571 <u>217478</u>	23 00 49.1923	+21 22 53.4144	4.81	К0	HD 217478	,Star,IR,*,	0.965	0.002	0.33	0.542	2.374	0.506
13	1.623	23 01 22.3224	+21 55 09.0300	5.65	F0	BD+21 4867	,Star,*,IR,	, 0.986	6.804E-4	9.572	0.334	2.489	0.317
14	1.677	23 01 15.5138	+20 18 00.1152	6.365	G5	BD+19 5044	,Star,*,IR,	, 0.992	2 3.557E-4	0.673	0.254	2.26	0.238
15	1.714 <u>219311</u>	23 14 44.6614	+20 53 13.9200	7.116	К0	HD 219311	,PM*,PM*,*,IR,	0.996	5 1.727E-4	13.217	0.18	2.169	0.168
16	1.731 <u>217385</u>	23 00 03.7903	+21 13 10.5816	6.662	F2	HD 217385	,Star,*,IR,	, 0.990	5 2.151E-4	0.539	0.188	2.476	0.178

Se	a	rcl	hC	a)	[c 1]

			-2) Search Cal Baramaters				
			3) SearchCal Parameters				
::	Q~ HR 8799	8]	Min. Magnitude (K) :	3.24			
]:	23 07 28.7156905667		Max. Magnitude (K) :	7.24			
]:	+21 08 03.302133882		Scenario :	0	Bright	🔵 Faint	
):	5.24		RA Range [mn] :	240.0			
			DEC Range [deg] :	30.0			
					Get	Calibrators)
							_

http://www.jmmc.fr/searchcal_page.htm







JMMC/SearchCal: developed for long-baseline interferometry

														_
13	1.623	23 01 22.3224	+21 55 09.0300	5.65	F0	BD+21 4867	,Star,*,IR,	0.986	6.804E-4	9.572	0.334	2.489	0.317	
14	1.677	23 01 15.5138	+20 18 00.1152	6.365	G5	BD+19 5044	,Star,*,IR,	0.992	3.557E-4	0.673	0.254	2.26	0.238	
15	1.714 <u>219311</u>	23 14 44.6614	+20 53 13.9200	7.116	К0	HD 219311	,PM*,PM*,*,IR,	0.996	1.727E-4	13.217	0.18	2.169	0.168	
16	1.731 <u>217385</u>	23 00 03.7903	+21 13 10.5816	6.662	F2	HD 217385	,Star,*,IR,	0.996	2.151E-4	0.539	0.188	2.476	0.178	
17	1.783 <u>219292</u>	23 14 30.5165	+20 26 32.1216	4.53	К2	HD 219292	,Star,**,*,IR,	0.958	0.008	0.477	0.597	8.987	0.554	
18	1.915	23 15 00.9744	+21 54 21.2436	6.42	К0	BD+21 4902	,Star,*,IR,	0.99	4.314E-4	0.781	0.282	2.216	0.263	
19	2.038	23 16 12.4874	+21 03 20.8476	3.699	K5	HD 219525	,Star,IR,*,	0.876	0.026	0.04	1.046	10.649	0.964	
20	2.084	22 58 50.7281	+20 36 28.9692	6.856	K2	<u>BD+19 5039a</u>	,Star,*,IR,	0.994	3.054E-4	1.766	0.228	2.409	0.211	
21	2.217	22 58 07.9622	+20 44 33.3204	5.771	К0	HD 217116	,Star,**,*,IR,	0.986	6.538E-4	0.429	0.347	2.23	0.324	
22	2.286	23 17 08.0170	+20 45 22.6656	6.964	G0	<u>BD+19 5081</u>	,Star,*,**,	0.996	2.083E-4	0.617	0.189	2.373	0.178	
23	2.394	22 57 12.7490	+21 06 32.7024	7.167	G5	BD+20 5251	,Star,*,IR,	0.996	1.677E-4	0.781	0.174	2.245	0.163	
						0								_
Filters —														
	<u> </u>			<u> </u>				DECC		1.0				_
🔽 Reje	ct stars farther thar	Maximum RA S	Separation (mn) : 6	0.0			Maximum	DEC Sepai	ration (degree	:): 1.0				
Reject stars with magnitude : below : 0.0														
Reject Spectral Types (and unknowns) :											5			
🗌 Reject Luminosity Classes (and unknowns) : 👘 🗌 🛛 V 🗹 V 🗹														
Reject Visiblity below : vis2 : 0.5														
Reject Visibility Accuracy above (or unknown): vis2Err/vis2 (%): 2.0														
Reje	Reject Variability													
🔽 Reje	Reject Multiplicity													
🔽 Reje	✓ Reject Invalid Object Types													
🗌 🗌 Diar	Diameter quality : Maximum chi square : 2.0 Maximum relative error (%) : 10.0													
search	Rearching calibrators done.													

http://www.jmmc.fr/searchcal_page.htm







In [4]: s1.separation(s2) Out[4]: <Angle 4.72223611 deg> In [6]: s1.separation(s3) Out[6]: <Angle 1.24333427 deg> In [7]: sep=s1.separation(s3) In [8]: sep_arcsec **Out[8]:** 4476.003380925846

- In [1]: from astropy import coordinates, units as u, wcs In [2]: s1 = coordinates.SkyCoord.from_name('HR8799') In [3]: s2 = coordinates.SkyCoord.from_name('HD220657')
- In [5]: s3 = coordinates.SkyCoord.from_name('HD218261')



Coronagraphy Visibility Tool (CVT)

selsin.





- **Open the Coronagraph Visibility Tool**
- On your laptop if you have it installed (command line or OSX binary)
- Run it on the Virtual Desktop on https://jwst-masterclass.science.stsci.edu otherwise
- Enter HR8799 "Search" to resolve it with Simbad, then press "update plots"
- Enter the b companion: PA=45° and separation=1.7" and press "update plots"
 - How many days per year the star is observable with JWST? •
 - When (convert to MM-DD) is it ideal to observe HR8799b with the LW bar? •
 - How much is the maximum roll angle around that date? **•**

Exercise: Search for a more efficient reference star, 3 methods







<u>https://jwst-masterclass.science.stsci.edu</u>



May take a few minutes to load

/spawn/jgirard@stsci.edu?next=%2Fhub%2F	♀ ⊌ ☆		
		jgirard@stsci.edu	•

Spawner Options

Spawn







Other



https://jwstmasterclass.science <u>stsci.edu</u>



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Jupyter Hub: virtual desktop for the CVT GUI



https://jwstmasterclass.scien <u>stsci.edu</u>

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jirard-40stsci-2eedu:1 (jovyan)					Ser	nd Ctr	IAItDel
					C	04:21	jovyan
launch CVT							

С	e	

STScI | SPACE TELESCOPE SCIENCE INSTITUTE

Jupyter Hub: virtual desktop for CVT, it runs!





https://jwstmasterclass.science. stsci.edu







Example Science Program: the HR8799 4-planet system



discovered by Marois, here Currie et al. 2014

•										
Та	rget Location									
SIM	BAD Target Reso	lver								
hr8	799	Searc								
ID:	HD 218396									
RA:	346.8696479	° (decin								
Dec	21.13425138	o (decim								
Eclip	tic coordinates:									
(I, b) = (356.7876°,	24.52479)								
Co	mpanions									
	PA (°)	Sep (")								
	45	1.7								
	325	1								
	0.00	0.00								
	Instrument	t								
NR	Cam Channel A	×								
Mask										
NRCA5_MASKLWB										
Aperture PA V3 PA										
Da	Date and Sampling									
Time	esteps per year:	360								
Rolls	s checked:	20								
	Update I	Plot								
_										
	Zoom to	fit								

Coronagraph Visibility Tool (CVT)

JWST Coronagraph Visibility Tool



+ Q ≒ 🖺





Jupyter Hub: best time to observe HR8799 b?









Jupyter Hub: roll angle "stroke"?

356

Day of year (from Jan 1)

358

354



JWST Coronagraph Visibility Tool









ETC Calculations







https://jwst.etc.stsci.edu/

Welcome to the JWST Exposure Time Calculator

Login

Work Anonymously

This release features new instrument modes, accuracy improvements, usability enhancements, and more: see the Release Notes for details, and be sure to review the Known Issues for this



ETC for Coronagraphy: PSF subtraction from a reference star



Instrument Filter/Disperser:	f335m/null
Extraction Apenture Position (arcsec):	[1.70, 0.00]
Wavelength of Interest used to Daloulate Scalar Values (microna):	3.35
Size of Extraction Aperture (areace):	0.08
Total Time Required for Strategy (seconds):	716.16
Total Exposure Time (accords):	358.08
Extracted Flux (o-/acc):	76.73
Standard Deviction in Extracted Flux (e-/sec):	0.82
Extracted Signal-to-Noise ratio:	93.65
Input Background Surface Brightness (MJy/sr):	0.14
Total Background Flux in Extraction Aperture (e-/sec):	00.36
Total Sky Background Flux in Extraction Aperture (e-/sec):	1.20
Fraction of Total Background due to Signal From Scene:	0.98
Average Number of Cosmic Rays per Pamp:	4.0e-3
Radius at which Contrast is Measured (arcsec):	1.00
Azimuth at which Contrast is Measures (degrees):	0.0
Contrast:	NaN

<u>jwst.etc.stsci.edu</u>



ETC Workbook for HR8799 b c d e: #27313

Open the Exposure Time Calculator (Google Search "JWST ETC" or jwst.etc.stsci.edu)

Make a Copy of the shared workbook.



Exposu	ire T	ime Calculator Edit	t - Expan	d 👻				Julien Girard -	Help 👻
Workb	ook ID	: 27040 Ma	aster Class: HR (8799bcde			NIRCam and MIRI Coronagraphic Imaging of the HR 8799 planetary system		
Calculations Scenes and Sources Upload Spectra Caveats and Limitations MIRI ▼ NIRCam ▼ NIRSS ▼ NIRSpec ▼ Image: Scene ★ Backgrounds Instrument Setup Detector Setup Strategy ID ▲ O Mode - λ - Scn - (s) - SNR - Δ 5 inircam coronagraphy 2.50 1 89.88 94.08 O Observation Scene ★ Scene ★ Scene ★ Backgrounds Instrument Setup Detector Setup Strategy									
MIRI	•][NIRCam - NIRISS -	NIRSpec -	0			Scene ★ Backgrounds Instrument Setup Detector Setup Strategy		
ID 🔺	Ø	Mode - λ -	- Scn -	(s) -	SNR - 🛛		Coronagraphy		
5		nircam coronagraphy 2.5	50 1	89.88	94.08		()))))))))))))))))))		
4		nircam coronagraphy 4.2	28 1	90.97	216.32		Observation Extraction		
3		nircam coronagraphy 3.5	56 1	89.88	218.87		Scene rotation		
2		nircam coronagraphy 2.9	99 1	89.88	188.75		0 deg ccw		
1	$[\checkmark]$	nircam coronagraphy 2.5	50 1	89.88	-102.42		PSF subtraction source		
-	-		-				2: Reference ▲ Optimal (PSF Autoscaling) Optimal (No PSF Autoscaling) Unsubtracted Science Scene ✓ PSF Subtraction Source Only		
							Calculation selected: 1, Mode: nircam coronagraphy	Reset Ca	alculate

ETC Workbook: PSF Subtraction Source (4 options)



ETC Workbook for HR8799 b c d e: #27313, downloaded files

	💽 Downloads			
$\langle \rangle$		* *	Q s	Search
Favorites	Name	Size	Kind	Date Added V
😭 igirard	wb27040_c1_2019-11-07_23.43.15_PSFsource		Folder	Today at 6:44 PM
	wb27040_c1_2019-11-07_23.43.15.tar	9.4 MB	tar archive	Today at 6:44 PM
Downloads	wb27040_c1_2019-11-07_23.21.44_UnsubtractedScene		Folder	Today at 6:43 PM
ScreenCaptures	b wb27040_c1_2019-11-07_23.21.44.tar	9.4 MB	tar archive	Today at 6:22 PM
	wb27040_c1_2019-11-07_23.20.50_Optimal_no-Autoscalling		Folder	Today at 6:22 PM
	b wb27040_c1_2019-11-07_23.20.50.tar	9.4 MB	tar archive	Today at 6:21 PM
WFIRST	wb27040_c1_2019-10-31_21.55.24_Optimal_Autoscalling		Folder	Today at 6:18 PM
🚞 JWST	The second secon		Folder	Today at 6:18 PM
MIRCam	Ineplot_wave_calc.fits	9 KB	FlexiblSystem	Today at 6:18 PM
	lineplot_extracted_flux_plus_bg.fits	9 KB	FlexiblSystem	Today at 6:18 PM
CWG	lineplot_extracted_bg_only.fits	9 KB	FlexiblSystem	Today at 6:18 PM
DF PDF	lineplot_extracted_noise.fits	9 KB	FlexiblSystem	Today at 6:18 PM
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ETC Workbook for HR8799 b c d e: #27313, downloaded files





ETC: Finding the best exposure parameters

Selecting the optimal combination of readout pattern, ngroups, nints and nexps is a trade-off

- More frames decreases read noise
- Shorter groups increases data volume
- Longer groups increases the chance of a cosmic ray hit during the group
- Shorter integrations make ramp fits more uncertain in the presence of nonlinearity
- More dithered exposures decreases flat field errors (not currently modeled by ETC!)
- Patterns that skip a lot of frames have higher read noise, but have slightly better duty cycle

ratio (SNR)

Signal-to-Noise

10¹



APT & smart (time) accounting

and suit









Home > Scientific Community > Software

What is APT?

The Astronomer's Proposal Tool (APT) is used to write, validate, and submit proposals for the Hubble Space Telescope and the James Webb Space Telescope.

Download and Installation Instructions

Linux

Mac OSX

www.stsci.edu/scientific-community/software/ astronomers-proposal-tool-apt

Q SEARCH

MENU

Windows

Current Release: 27.3

Released: September 16, 2019

This upgrade is not required for HST Proposers.

This upgrade is recommended for people working on JWST programs. Read more 🗹

Previous Release Information



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Changing attitude

- 1. Update observatory pointing and roll
- 2. Let disturbances settle
- 3. Reacquire guide star
 - + Fine guide (always)
 - +Track (>0.06")
 - +Acquisition (>25")
 - Identification (new visit)



Reference star & overheads: slew, settle, reacquire guide star

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M2V	1.12	1.11	1.10	1.10	1.08	1.08	1.07	1.06	1.06	1.06	1.05	1.05	1.04	1.00	0.93	0.75	0.39	0.00	0.02	1.05	1.03	1.02	0.99	0.92
M0V	0.69	0.68	0.66	0.66	0.64	0.64	0.63	0.62	0.61	0.60	0.60	0.59	0.58	0.53	0.44	0.23	0.00	0.40	0.50	0.60	0.58	0.55	0.50	0.42
K7V	0.32	0.32	0.30	0.30	0.28	0.27	0.26	0.25	0.24	0.23	0.23	0.22	0.21	0.16	0.09	0.00	0.23	0.75	0.82	0.23	0.20	0.18	0.15	0.08
K5V	0.12	0.12	0.11	0.10	0.09	0.08	0.07	0.07	0.06	0.06	0.05	0.05	0.04	0.02	0.00	0.08	0.45	0.92	0.99	0.05	0.04	0.03	0.01	0.01
K2V	0.06	0.06	0.05	0.05	0.03	0.03	0.03	0.02	0.01	0.01	0.01	0.01	0.01	0.00	0.01	0.16	0.53	0.99	1.05	0.01	0.01	0.00	0.01	0.03
K0V	0.03	0.03	0.02	0.02	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.01	0.04	0.21	0.58	1.03	1.09	0.00	0.00	0.00	0.02	0.05
G8V	0.02	0.02	0.02	0.02	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.04	0.22	0.59	1.04	1.10	0.00	0.00	0.00	0.02	0.06
G5V	0.02	0.02	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.05	0.23	0.60	1.04	1.10	0.00	0.00	0.01	0.02	0.06
G2V	0.02	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.05	0.24	0.61	1.05	1.11	0.00	0.00	0.01	0.02	0.07
G0V	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.06	0.24	0.62	1.05	1.11	0.00	0.00	0.01	0.02	0.07
F8V	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.06	0.25	0.62	1.06	1.12	0.00	0.00	0.01	0.03	0.08
F5V	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.07	0.26	0.63	1.06	1.12	0.00	0.00	0.01	0.03	0.08
F2V	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.03	0.08	0.27	0.64	1.07	1.13	0.00	0.01	0.02	0.04	0.09
F0V	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.03	0.08	0.28	0.65	1.08	1.14	0.01	0.01	0.02	0.04	0.10
A5V	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.04	0.10	0.30	0.67	1.09	1.15	0.01	0.02	0.03	0.06	0.11
A3V	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.02	0.04	0.10	0.30	0.67	1.09	1.15	0.01	0.02	0.03	0.05	0.11
A1V	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.02	0.05	0.12	0.32	0.68	1.11	1.16	0.02	0.03	0.04	0.07	0.14
A0V	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.02	0.02	0.03	0.06	0.12	0.32	0.69	1.11	1.17	0.02	0.03	0.04	0.07	0.14

Reference Star SpType

credit: Jarron Leisenring (pyNRC) PACE TELESCOPE



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APT: special requirements, Non-interruptible sequence

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APT: special requirements, Non-interruptible sequence



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APT: special requirements, Fiducial Pointing Override

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APT: special requirements, Fiducial Pointing Override

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APT: PSF reference stars



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APT: PSF reference stars



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A JWST Approved Proposal 3 (Unsaved)	HR 8799 bcde - NIRCam - Roll 2 - MASKLWB (Obs 5) of JWST Approved Proposal 1194 (Unsaved)
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Ø Proposal Information	Number 5 Status: IMPLEMENTATION
Targets	Label HR 8799 bcde - NIRCam - Roll 2 - MASKLWB
Governations	Instrument NIRCAM C
 See HK 8799 DCce HR 8799 - NIRCam - Roll 1 - MASK430R (Obs 1) 	Template NIRCam Coronagraphic Imaging
HR 8799 bcde - NIRCam - Roll 1 - MASKLWB (Obs 2)	
Ref star - NIRCam - MASK430R (Obs 3)	larget THR8799
Ref star - NIRCam - MASKLWB (Obs 4)	Splitting Distance Number of Visits
HR 8799 bcde - NIRCam - Roll 2 - MASKLWB (Obs 5) MASKLWB (Obs 6)	Science Total Charged
HR8799 1065C (Obs 7)	Duration (secs) 6663 13071
HR8799 1140C (Obs 8)	Data Volume 980 MB
HR8799 1550C (Obs 9)	
HR8799 2300C (Obs 10)	NIRCam Coronagraphic Imaging 🔥 Special Requirements Comments
REF 1055C (Obs 11)	Module A 🗘
REF 1140C (Obs 12) REF 1550C (Obs 13)	Coronagraphic Mask MASKLWB
REF 2300C (Obs 14)	Target Acquisition Parameters
P Observation Links	Acq Target Acq Tilter Acq Target Brightness
	Target ACQ Same Target as Observation 😋 🖒 F335M 🗘 BRIGHT (ND Square) 😋
	Acq Readout Pattern Acq Groups/Int Acq Integrations/Exp Acq Total Integrations Acq Total Exposure Time Acq ETC Wkbk.Calc ID ETC
	Astrometric Confirmation Image Parameters
	Obtain Astrometric Confirmation Images? • Yes · No
	Conf. Readout Pattern Conf. Groups/Int Conf. Integrations/Exp Conf. Total Dithers Conf. Total Integrations Conf. Total Exposure Time
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	Subarray SUB320
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APT: smart accounting, astrometric confirmation image (TA)

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Target Acquisition	n Parameters						
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Subarray	SUB320						
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APT: optional, full frame astrometric confirmation images (TA)





Proposal ID	1194	STScI Edit Number	r 7
Category	сто ᅌ		
Pure Parallel Proposal			
Cycle	1		
	Explain unsched	ulable observations	
X Charged Time (hours)	23.60	Charg	ed Time
Allocated Time (hours)	23.40		
Proposal Size	SMALL		
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STScI Only PPS DB Overrid	es		
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	Edit PPS D	B Overrides	Overrides: None

with Confirmation Images



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APT: smart accounting, astrometric confirmation image (TA)

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APT: optional, full frame astrometric confirmation images (TA)

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APT: smart accounting, astrometric confirmation image (TA)

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O Proposal Information of JWST Approved Proposal 1194 (Unsaved)		
HR 8799 planetary system and planet search		
n, European MIRI, and Telescope Team GTO scientists will execute a series of coronagraphic measurer are two-fold. viously unknown planets using NIRCam in the F322W2 and F444W filters with the round 430 mask be consitivity to masses less than 1 Mlup at F444W and will use F322W2 to reject background stars and	nents using NIRCam and ing used for both filters.	i MIRI. The This
sensitivity to masses less than 1 Mjup at F444W and Will use F322W2 to reject background stars and program is the physical characterization of the known planets, HR8789bcde, using NIRCam and MIRI cercise an engineering mode of the telescope to aggressively push the Inner Working Angle and detect separated planets. Six medium-band filters will be used in conjunction with the long wavelength bar. leg) and a reference star to assist with suppression of residuals in the coronagraphic image. The MIRI of the MIRI 4 Quadrant Phase mask (4QPM) coronagraph.	galaxies. multi-filter photometry t HR8799e while also me The NIRCam observatior team will observe the sy	: The asuring is will use /stem
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Proposal ID	1194	STScI Edit Numbe	er 7
Category	GTO 🗘		
Pure Parallel Proposal			
Cycle	1		
	Explain unsched	ulable observations	
Science Time (hours)	11.60		
X Charged Time (hours)	23.45	Char	ged I me
Allocated Time (hours)	23.40		
Proposal Size	SMALL		
Allow Restricted	(this sessi	on only)	
STScI Only PPS DB Overrid	es		
	Edit PPS D	B Overrides	Overrides: None

Taking astrometric confirmation images add 0.15h (9 min) in this case

APT: smart accounting, astrometric confirmation images (TA)

with no Astrometic Confirmation Images



APT: smart accounting, closer PSF reference star



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99 planetary system and planet search	ants a	usina	NIRCam and	MIDI	The
-fold. Unknown planets using NIRCam in the F322W2 and F444W filters with the round 430 mask bein ivity to masses less than 1 MJup at F444W and will use F322W2 to reject background stars and gr am is the physical characterization of the known planets, HR8789bcde, using NIRCam and MIRI n an engineering mode of the telescope to aggressively push the Inner Working Angle and detect H ed planets. Six medium-band filters will be used in conjunction with the long wavelength bar. Th d a reference star to assist with suppression of residuals in the coronagraphic image. The MIRI to MIRI 4 Quadrant Phase mask (4QPM) coronagraph.	g use alaxie nulti- HR87 he NI sam v	d for es. filter 99e w RCam vill ob	both filters. photometry hile also me observation serve the sy	This . The easuring is will u	J
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8799 planetary system and planet search	
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isly unknown planets using NIRCam in the F322W2 and F444W filters with the round 430 mask be sitivity to masses less than 1 MJup at F444W and will use F322W2 to reject background stars and ogram is the physical characterization of the known planets, HR8789bcde, using NIRCam and MIR ise an engineering mode of the telescope to aggressively push the Inner Working Angle and detect arated planets. Six medium-band filters will be used in conjunction with the long wavelength bar, and a reference star to assist with suppression of residuals in the coronagraphic image. The MIR he MIRI 4 Quadrant Phase mask (4QPM) coronagraph.	eing used for both filters. This d galaxies. RI multi-filter photometry. The ct HR8799e while also measuring The NIRCam observations will use RI team will observe the system
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	Astronomer's Proposal Tools Version 273 IWST PRD: PRDOPSSOC_M-025 _ IWST Approved Proposal 3 (Upsaved)	
	Astronomer s Proposar iouis version 27.5 GWST PRD. PRDOP350C-M-025 - GWST Approved Proposar 5 (Disaved)	
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JWST Approved Proposal 1194 (Unsaved)	Proposal Information of JWST Approved Proposal 3 (I	Unsaved)
JWST Approved Proposal 3 (Unsaved)		
Proposal Information	Title NIRCam Coronagraphy Example	
Targets	Abstract This program can be accessed as an example within APT demostration programs.	
Observations	It demonstrated a standard coroangraphic sequence with one science target (at two roll angles	and one PSF ref star, observed with two NIRCam coronagraph
Sequence Beta Pic Sequence	masks.	
NIRCam F210M Wedge, Koll 1 (Obs 1)	The 6-observation coronagraphic sequence is non-interruptible, and includes a roll dither after	er the first pair of observations. A roll range is used to constrain
NIRCam F210M Wedge Roll 2 (Obs 3)	the nominal angles.	
NIRCam F430M Sombrero, Roll 2 (Obs 4)	Values entered are legal, but not necessarily realistic or scientifically valid. This example is fo	or demonstration purposes only.
NIRCam F210M Wedge, PSF REF (Obs 5)	For a companion JWST User Support Documentation article describing this example, see the fo	llowing link:
NIRCam F430M Sombrero, PSF REF (Obs 6)		
P Observation Links	https://jwst-docs.stsci.edu/display/JPP/APT+Coronagraphic+Sequence+Examples	
	Updated and current for APT 25.4.2 (Jan. 2018).	
	Proposal ID 3 STScI Edit Number 4	
	Category GO 🗘 Calibration Treasury	
	Pure Parallel Proposal	
	Cycle 1	
	Explain unschedulable observations	
	Science Time (nours) 9.82	unting
	A Charged Time (hours) 18.19 Run Smart Accounting	unung
	Data volume (MB) 4629.39	
	Allocated Time (hours) 25.00	
	Proposal Size SMALL	
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APT: smart accounting, astrometric confirmation image (TA)

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APT: smart accounting, astrometric confirmation image (TA)

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tandard coroangraphic sequence with one science target (at two roll angles) and one PSF ref star, observed with two NIRCam coronagraph
coronagraphic sequence is non-interruptible, and includes a roll dither after the first pair of observations. A roll range is used to constrain
egal, but not necessarily realistic or scientifically valid. THis example is for demonstration purposes only.
ST User Support Documentation article describing this example, see the following link:
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Extra Slides & Extra Resources

jwsthelp.stsci.edu

James Webb Help Desk

Your JWST gateway. Report issues and submit requests.

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	APT Support Request assistance with the Astronomer's Proposal Tool (APT)	Constraints & Schedulability Ask questions about schedulability and observing with JWST	Coronagraphy Ask about NII MIRI coronag imaging
	view Details	View Details	view Details
	Data Analysis Tools for JWST Request assistance with STScI-developed data analysis to ba.	ETC Support Request assistance with the Exposure Time Calculator (ETC)	JWST Science Policies Request assis Science Polic
	View Plans	View Details	View Details
Knowledge	JWST SN Requests & Issues Submit JWST Requests and Issues related to ServiceNow	MIRI Support Request assistance with the Mid-Infrared Instrument (MIRI)	NIRCam Support Request assis with the Near Camera (NIRC
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Q	NIRISS Support Request assistance with the Near-Infrared Imager and Slitless Spectrograph (NIRISS)	NIRSpec Support Request assistance with the Near-Infrared Spectrograph (NIPSper)	Office of Public Outrea Contact the S Office of Public Outreach abo
	View Details	View Details	View Details
	Pipeline Support Request assistance with the JWST pipeline View Details	Solar System Observing Ask questions about proposal writing for solar system targets with IWCT View Details	Time-Series Observation Request assis making time- observations transiting aver View Details
phy tion strategy	WebbPSF / JWST Telescope Request assistance with the WebbPSF tool	JWST General Support Request general JWST support for issues not	MAST Archive Support Request gene Archive support
	ontical system	covered by another	issues not cov another cated
screenshots.	View Details	View Details	View Details

Proposal Planning Workshop: material, presentations

🛞 STScI

NASA's James Webb Space Telescope

Developed in partnership with ESA and CSA. Operated by AURA's Space Telescope Science Institute

JWST SCIENCE 🗮	NEWS & EVENTS 📃	INSTRUMENTATION
DOCUMENTATION		
NEWS & EVENTS > Eve	ents	

Past Events

Planning Solar System Observations with JWST - ESTEC venue

Science Meeting

December 13 - 15, 2017

Noordwijk, Netherlands ESTEC

This 2.5-day workshop will include a mixture of presentations about the promise of JWST for solar system science, specifics on observer planning tools and observatory capabilities, and hands-on training and Q&A with the planning tools. Observations of solar system targets approved for guaranteed-time observers (GTOs) and through the Early Release Science (ERS) program will be summarized. The workshop...

JWST Proposal Planning Workshop

Training Workshop • December 11 - 14, 2017 • Caltech, Pasadena, CA This workshop will take place shortly after the announcement of the programs selected under the first JWST open call for proposals (the Directory Discretionary Early Release Science Programs), and shortly before their observing files (meant to serve as models for the general observer community) become public. Therefore, the workshop will coincide with active proposal preparation for the next open...

https://jwst.stsci.edu/events

JISCI SPACE TELESCOPE SCIENCE INSTITUTE

Select a collection MAST Observations by Object Name of	r RA/Dec 🎽		and enter target: HIP 65426 Search				đ			
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Approved programs on MAST: example of ERS #1386 (Hinkley)

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